

Variational formulations of General Relativity

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Interdisciplinary junior scientist workshop:
Mathematical General Relativity

Introduction

Variational principles in General Relativity:

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- metric picture:

$$\mathcal{L}_g = \mathcal{L}_g (\mathfrak{g}_{\mu\nu}, \mathfrak{g}_{\mu\nu,\alpha}, \mathfrak{g}_{\mu\nu,\alpha\beta}, \phi, \phi_{,\nu}) = \mathcal{L}_H(\overset{\circ}{R}) + \mathcal{L}_{\text{matt}},$$

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- affine picture:

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Idea:

All of these formulations are equivalent *on shell*.

Metric Hilbert Lagrangian:

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$$\mathcal{L}_H(\mathring{R}) = \frac{\sqrt{|\det g|}}{16\pi} g^{\mu\nu} \mathring{R}_{\mu\nu} := \pi^{\mu\nu} \mathring{R}_{\mu\nu},$$

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$$\frac{1}{8\pi} \overset{\circ}{G}^{\mu\nu} = \phi^{;\mu} \phi^{;\nu} - \frac{1}{2} g^{\mu\nu} (\phi_{,\alpha} \phi^{,\alpha} + m^2 \phi^2),$$

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Important equality:

$$\begin{aligned} \mathcal{P}^{\lambda\mu}{}_\kappa \delta \overset{\circ}{\Gamma}^\kappa{}_{\lambda\mu} + \partial_\nu (p^\nu \delta \phi) &= \left(\mathcal{P}^{\lambda\mu}{}_\kappa - \frac{\partial \mathcal{L}_{\text{matt}}}{\partial \overset{\circ}{\Gamma}^\kappa{}_{\lambda\mu}} \right) \delta \overset{\circ}{\Gamma}^\kappa{}_{\lambda\mu} + \\ &+ \left(\overset{\circ}{\nabla}_\nu p^\nu \right) \delta \phi + p^\nu \delta \left(\overset{\circ}{\nabla}_\nu \phi \right), \end{aligned}$$

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$$\mathcal{L}_{\text{matt}} = -\frac{\sqrt{|\det g|}}{2} \left[\left(\overset{\circ}{\nabla}_{\nu} X^{\alpha} \right) \left(\overset{\circ}{\nabla}^{\nu} X_{\alpha} \right) + m^2 X_{\alpha} X^{\alpha} \right].$$

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Track to affine picture

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We have to transform the problematic term with a connection:

$$\mathcal{P}^{\lambda\mu}_{\kappa} \delta \overset{\circ}{\Gamma}^{\kappa}_{\lambda\mu} = \partial_{\kappa} (\mathcal{R}^{\mu\nu\kappa} \delta g_{\mu\nu}) - \left(\overset{\circ}{\nabla}_{\kappa} \mathcal{R}^{\mu\nu\kappa} \right) \delta g_{\mu\nu}.$$

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Variation of the metric Lagrangian:

$$\begin{aligned} \delta \mathcal{L}_g &= \left[\frac{\partial \mathcal{L}_{\text{matt}}}{\partial g_{\mu\nu}} - \frac{1}{16\pi} \overset{\circ}{G}^{\mu\nu} - \overset{\circ}{\nabla}_\kappa \mathcal{R}^{\mu\nu\kappa} \right] \delta g_{\mu\nu} + \\ &+ \partial_\kappa \left(\mathcal{R}^{\mu\nu\kappa} \delta g_{\mu\nu} + p^\kappa \delta \phi + \pi_\kappa^{\lambda\mu\nu} \delta \overset{\circ}{\Gamma}^\kappa_{\lambda\mu} \right). \end{aligned}$$

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Observation:

Metric appears in two ways: as a control parameter (δg) and momentum (π).

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Remark:

In affine picture metric appears as a momentum!

Transformation:

$$\partial_{\kappa} (\mathcal{R}^{\mu\nu\kappa} \delta g_{\mu\nu}) = \partial_{\nu} (\pi_{\kappa}^{\lambda\mu\nu} \delta N_{\lambda\mu}^{\kappa}) + \delta \left[\overset{\circ}{\nabla}_{\kappa} \mathcal{R}_{\sigma}^{\sigma\kappa} \right],$$

Transformation:

$$\begin{aligned}\partial_\kappa (\mathcal{R}^{\mu\nu\kappa} \delta g_{\mu\nu}) &= \partial_\nu (\pi_\kappa^{\lambda\mu\nu} \delta N_{\lambda\mu}^\kappa) + \delta \left[\overset{\circ}{\nabla}_\kappa \mathcal{R}_\sigma^{\sigma\kappa} \right], \\ N_{\lambda\mu}^\kappa &= \frac{16\pi}{\sqrt{|\det g|}} \left[\mathcal{R}_{\lambda\mu}{}^\kappa - \frac{1}{2} \mathcal{R}_\sigma^{\sigma\kappa} g_{\lambda\mu} + \right. \\ &\quad \left. - \frac{2}{3} \left(\delta_{(\lambda}^\kappa \mathcal{R}_{\mu)\sigma}{}^\sigma - \frac{1}{2} \mathcal{R}^\sigma_{\sigma(\lambda} \delta_{\mu)}^\kappa \right) \right], .\end{aligned}$$

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Result:

$$\delta \mathcal{L}_g = \partial_\kappa \left[p^\kappa \delta \phi + \pi_\kappa^{\lambda\mu\nu} \delta \left(\overset{\circ}{\Gamma}^\kappa_{\lambda\mu} + N_{\lambda\mu}^\kappa \right) \right] + \delta \left[\overset{\circ}{\nabla}_\kappa \mathcal{R}_\sigma^{\sigma\kappa} \right].$$

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We have an affine description of our theory!

Variation of the affine Lagrangian :

$$\begin{aligned}\delta\mathcal{L}_A &= \partial_\kappa \left[p^\kappa \delta\phi + \pi_\kappa^{\lambda\mu\nu} \delta\Gamma^\kappa_{\lambda\mu} \right] = \\ &= \partial_\kappa (p^\kappa \delta\phi) + \pi^{\mu\nu} \delta K_{\mu\nu} + \nabla_\nu \pi_\kappa^{\lambda\mu\nu} \delta\Gamma^\kappa_{\lambda\mu} .\end{aligned}$$

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Comments:

$K_{\mu\nu}$ is only the symmetric part of the Ricci tensor (for general symmetric connection there also exist a skew-symmetric part!)

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Vacuum with cosmological constant:

$$\mathcal{L}_g = \frac{\sqrt{|\det g|}}{16\pi} \left(\overset{\circ}{R} - 2\Lambda \right)$$

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If we want to take into account whole curvature we have to take the more general momentum of the connection, which symmetric part stays the same as previously.

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What is next?

I want to apply the last (traceless) part of Riemann curvature part and believe, that it could describe some kind of dark matter.

QUESTIONS?